

Note 365(5): Motion of stars in a Whirlpool Galaxy

In conventional classical dynamics the motion is governed by the Binet equation:

$$F = -\frac{L^2}{mr^3} \left( \frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) + \frac{1}{r} \right) \quad (1)$$

If:

$$r = \frac{r_0}{\theta} \quad (2)$$

Then:

$$\frac{d^2}{d\theta^2} \left( \frac{1}{r} \right) = 0 \quad (3)$$

So

$$F = -\frac{L^2}{mr^3} \quad (4)$$

and

$$m\ddot{r} = 0 \quad (5)$$

In this case:

$$\begin{aligned} \frac{dr}{dt} &= \frac{dr}{d\theta} \frac{d\theta}{dt} = \frac{L}{mr^2} \frac{dr}{d\theta} \\ &= -\frac{L}{mr^2} \frac{r_0}{\theta^2} = -\frac{L}{mr_0} \end{aligned} \quad (6)$$

So

$$r = -\frac{L}{mr_0} t \quad (7)$$

There is a contradiction between eqs. (4) and (7) because eq. (4) is an outward, centrifugal force and eq. (7) indicates that  $r$  becomes more negative with increasing  $t$ .

This contradiction is resolved by using:

$$\omega = \frac{d\theta}{dt} = -\frac{L}{mr} \quad (8)$$

2) i.e. by reversing the sign of the angular momentum:

$$\underline{L} = \underline{r} \times \underline{p} \quad - (9)$$

to

$$\underline{L} = -\underline{r} \times \underline{p} \quad - (10)$$

This is equivalent to reversing the motion of  $\underline{p}$  or  
time reversal:  $t \rightarrow -t$  - (11)

The magnitude of  $\underline{L}$ , denoted  $L$ , is not changed  
by motion reversal, so:

$$\hat{T}(L) = L \quad - (12)$$

where  $\hat{T}$  is the motion reversal operator. The angular  
velocity is

$$\underline{\omega} = \omega \underline{k} \quad - (13)$$

and

$$\hat{T}(\underline{\omega}) = -\underline{\omega} \quad - (14)$$

so eq. (8) results, QED.

Motion reversal changes eq. (7) to

$$r = \frac{L}{m r_0} t \quad - (15)$$

if

$$\frac{d\theta}{dt} = -\frac{L}{m r^2} \quad - (16)$$

The centrifugal force (4) has the property:

$$\hat{T}(F) = F \quad - (17)$$

because:

$$\hat{T}\left(\frac{L^2}{m r^3}\right) = \frac{L^2}{m r^3} \quad - (18)$$

Eqs. (4) and (15) mean that the stars

3) move outwards.

However, if we use:

$$r = -\frac{r_0}{\theta} \quad - (19)$$

$$\omega = -\frac{L}{mr^2} \quad - (20)$$

and

it follows that:

$$r = -\frac{L}{mr_0} t. \quad - (21)$$

The centrifugal force from eqs. (4) and (19) is:

$$F = \frac{L^2 \theta^3}{mr_0^3} \quad - (22)$$

and is positive valued. Eqs. (21) and (22) show  
near that the stars move inwards.

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